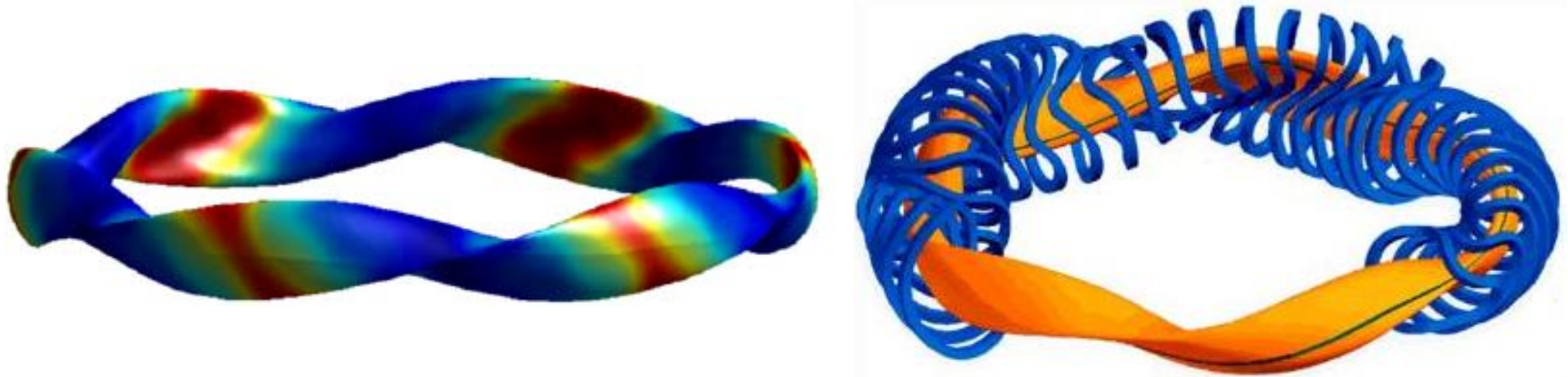


# Flow, current, & electric field in omnigenous stellarators



Matt Landreman

with Peter J Catto

MIT Plasma Science & Fusion Center

Oral 204 – Sherwood Fusion Theory Meeting

Tuesday May 3, 2011

*Supported by U.S. D.o.E.*



# Preview

W7-X can be – and any reactor must be – nearly omnigenous.

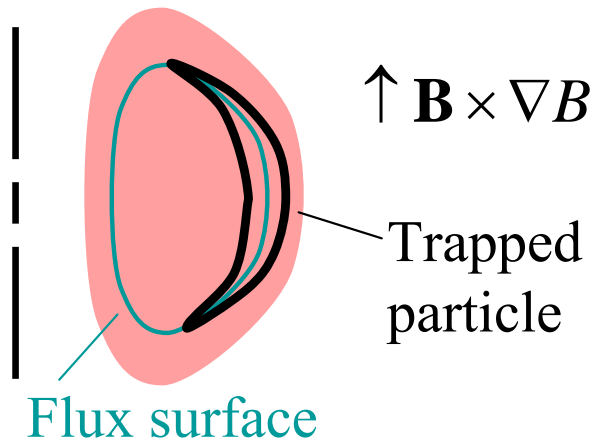
## Omnigenous stellarators:

- More general than quasisymmetric devices.
- **B** has nice properties.
- Formulae for current & flow simplify dramatically.
- Universal radial electric field.

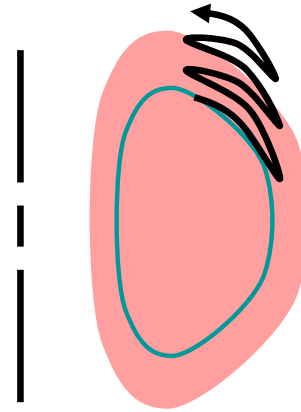
Landreman and Catto, PPCF **53**, 035106 (2011).

# Omnigenity = no unconfined orbits.

## Tokamak



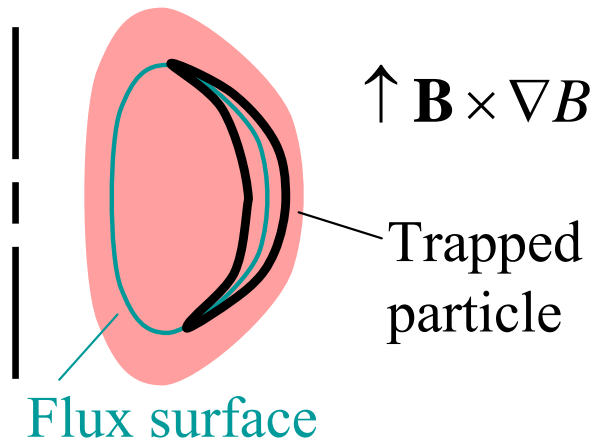
## Stellarator



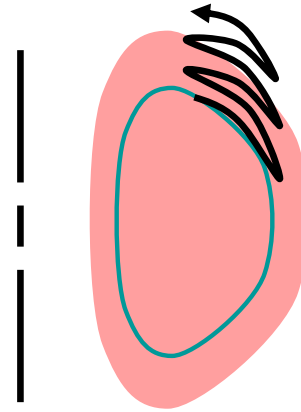
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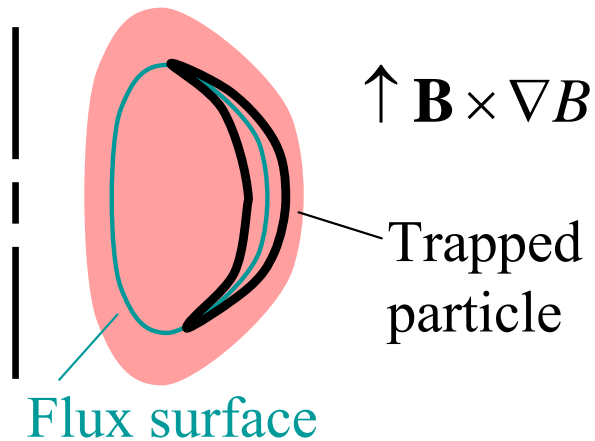
- Unconfined  $\alpha$  particles can damage plasma-facing components.

For a reactor, then, a stellarator must be nearly *omnigenous*:

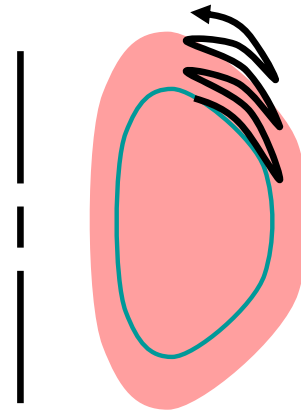
$$0 = \Delta \psi \text{ per bounce} = \oint_{\text{bounce}} (\mathbf{v}_d \cdot \nabla \psi) dt \quad \text{for all } \mu.$$

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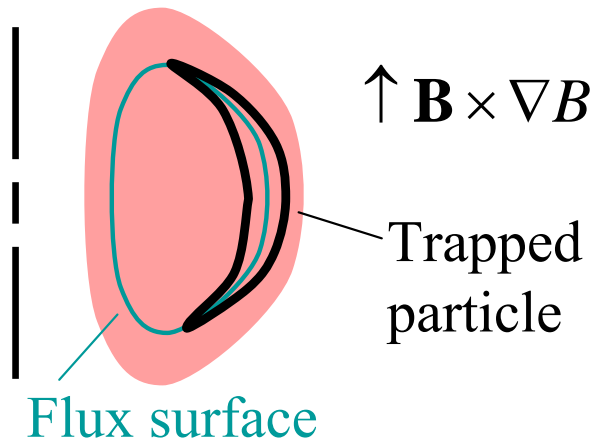
$J$  is a flux function,

where  $J = \oint v_{\parallel} dl$

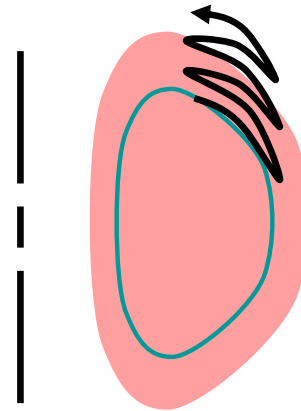
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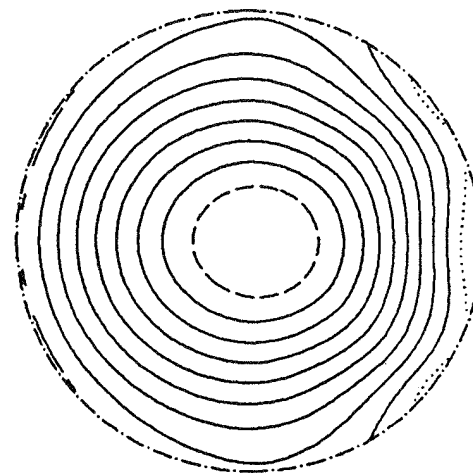
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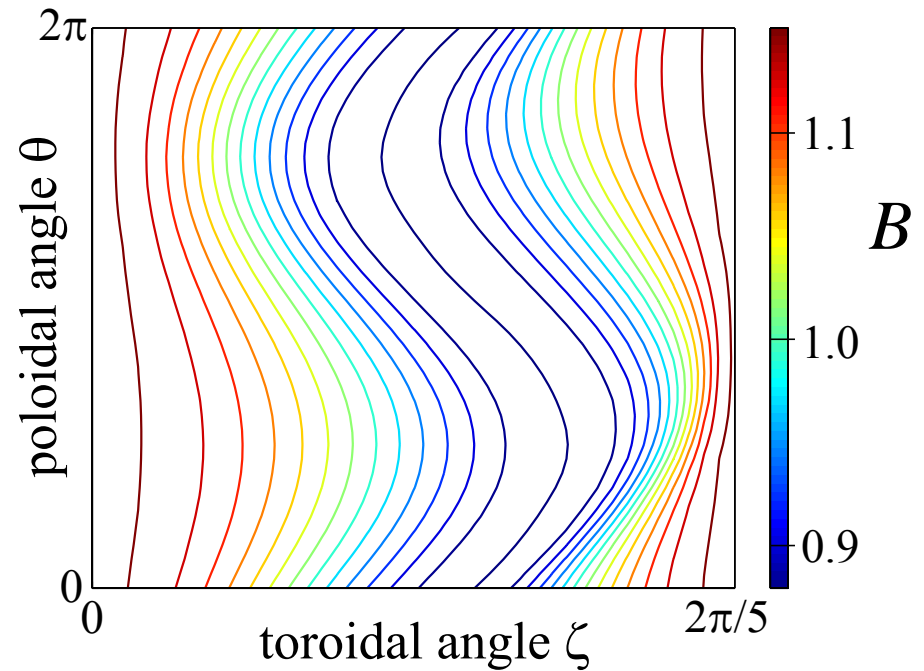
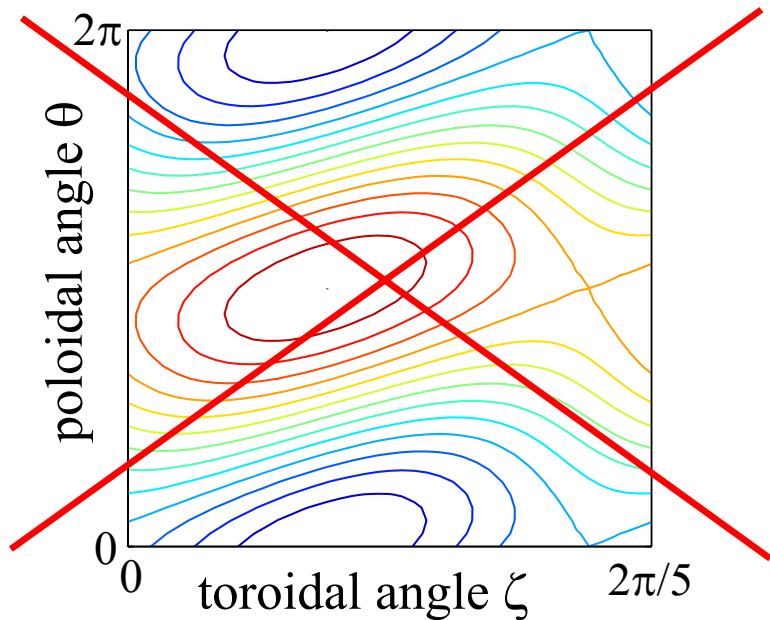


←  $J$  contours for W7-X  
 in  $(\sqrt{\psi}, \theta)$   
 polar coordinates

# Omnigenity places strong constraints on $B$ .

If  $\oint (\mathbf{v}_d \cdot \nabla \psi) dt = 0$ , then

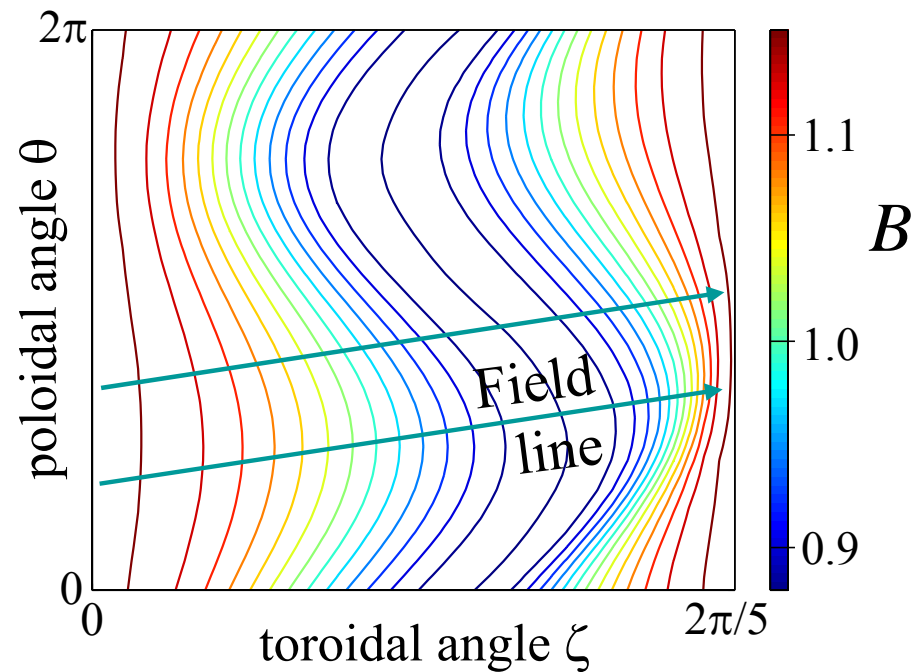
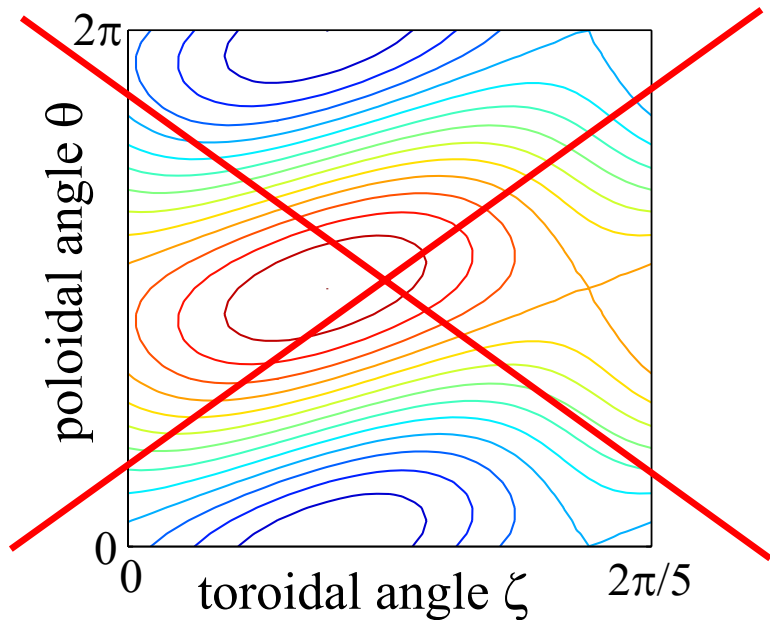
- All  $B$  contours link the torus toroidally, poloidally, or both.



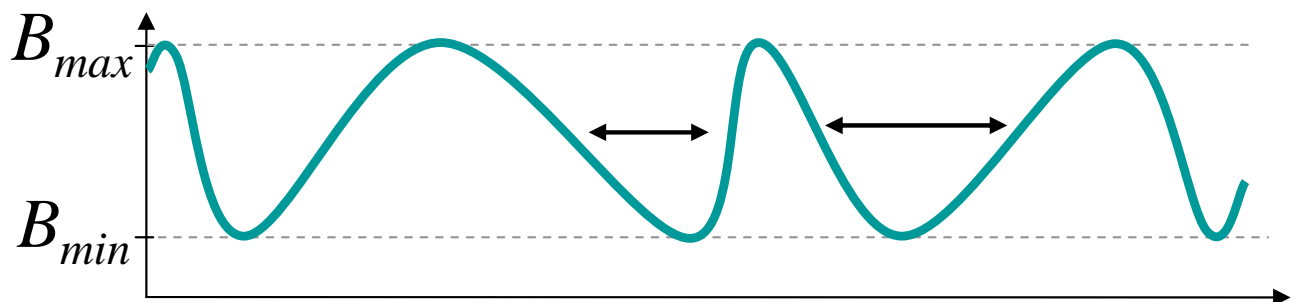
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If  $\oint (\mathbf{v}_d \cdot \nabla \psi) dt = 0$ , then

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- Each maximum & minimum of  $B$  along a field line is the same:





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Quasisymmetry:  $B = B(\psi, M\theta - N\zeta)$ .

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Neoclassical formulae simplify dramatically. Explicit non-numerical forms possible:

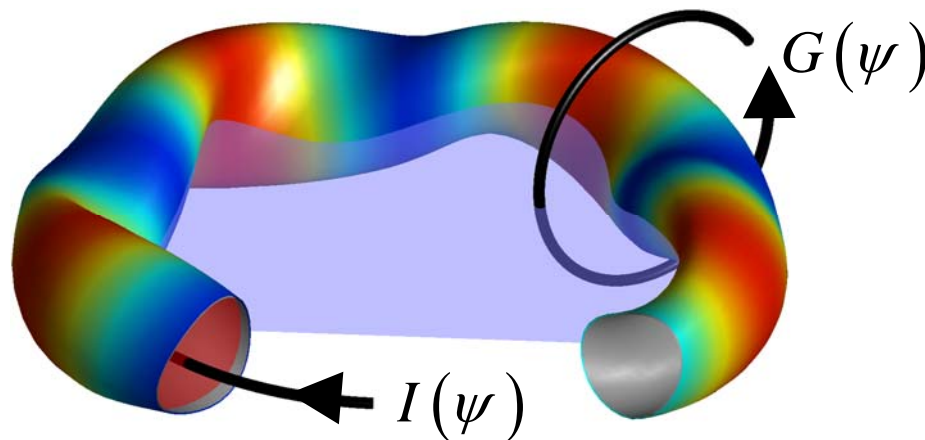
$$\text{Quasisymmetry: } \langle j_{\parallel} B \rangle = -4.8\sqrt{\varepsilon}q \left( \frac{dp_i}{d\psi} + \frac{dp_e}{d\psi} - 0.74n_e \frac{dT_e}{d\psi} - 1.17n_e \frac{dT_i}{d\psi} \right) \frac{MG + NI}{M - qN}$$

*Pytte & Boozer PoF (1981), Boozer PoF (1983)*

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where  $G(\psi)$  = poloidal current  
outside the flux surface,

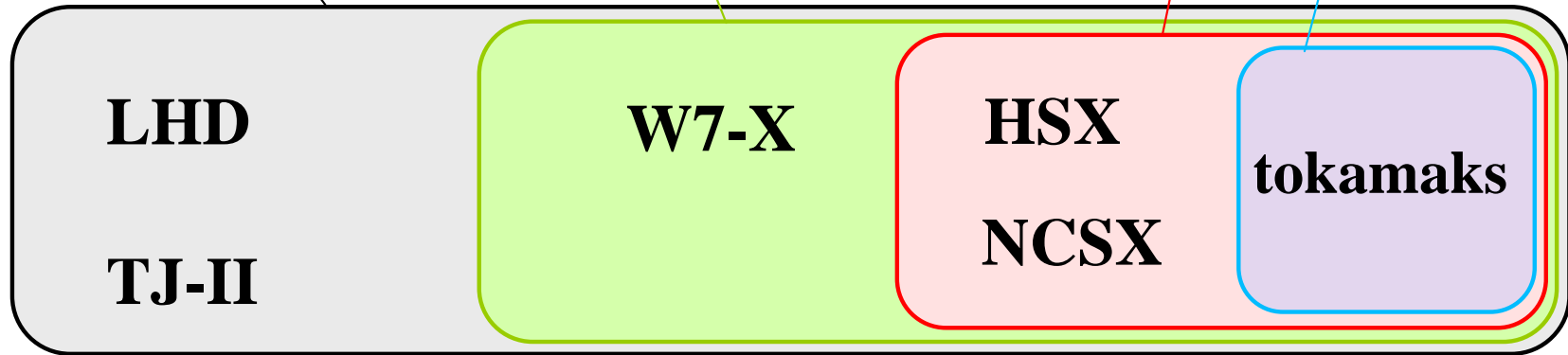
$I(\psi)$  = toroidal current  
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# Omnigenity is more general than quasisymmetry.

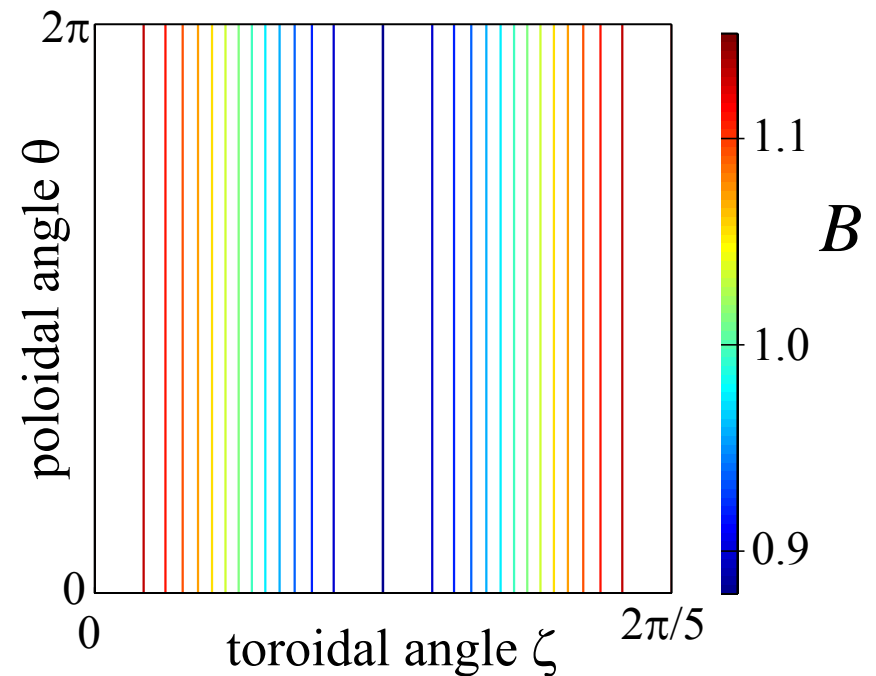
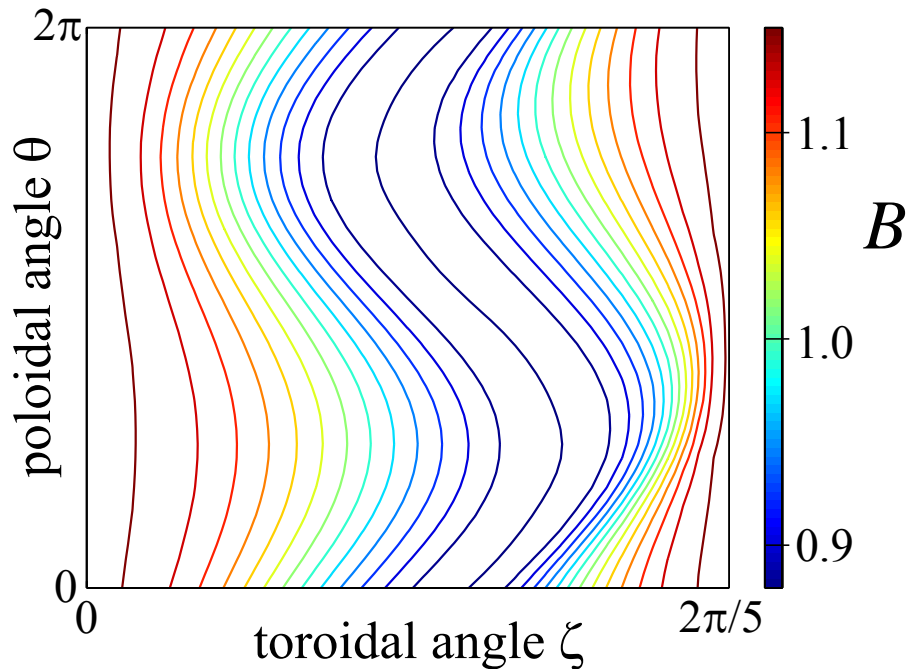
*Cary & Shasharina, PoP (1997), PRL (1997)*

All toroidal fields  $\supset$  Omnigenous  $\supset$  Quasisymmetric  $\supset$  Axisymmetric



Omnigenity:  $B$  contours may be curved

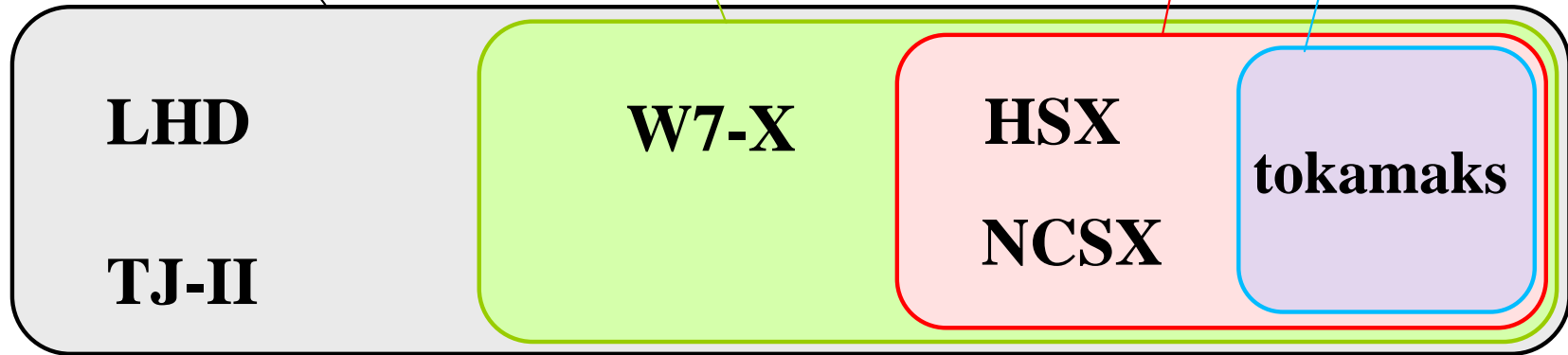
Quasisymmetry:  $B$  contours must be straight



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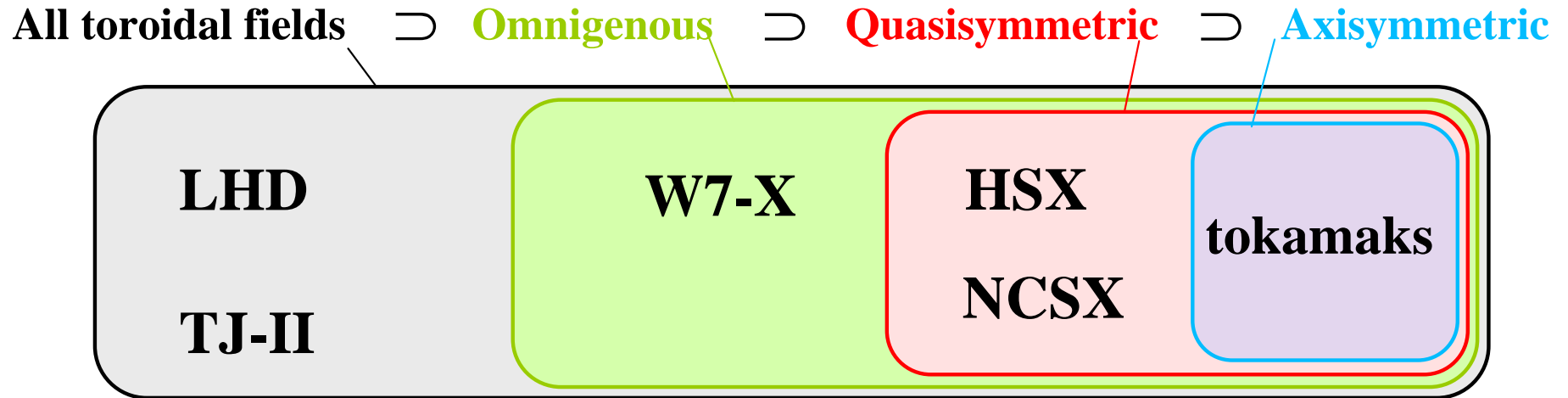
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Quasisymmetric fields **may** have reduced instability & turbulent transport:

- Flow speeds comparable to  $v_{th,i}$  are only permitted in quasisymmetric fields.  
(*Helander, PoP 2007*)
- Flows are clamped to a weakly sheared neoclassical value except in quasisymmetry, so quasisymmetric fields permit larger flow shear.  
(*Helander & Simakov, PRL 2008*)

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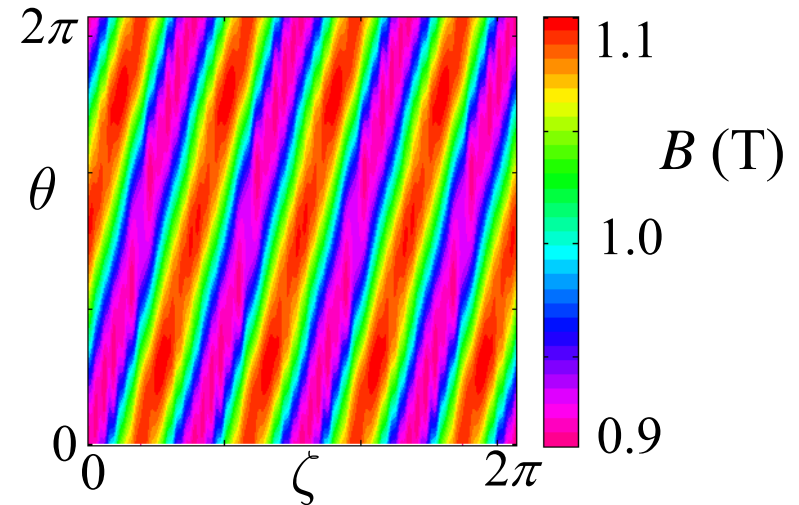
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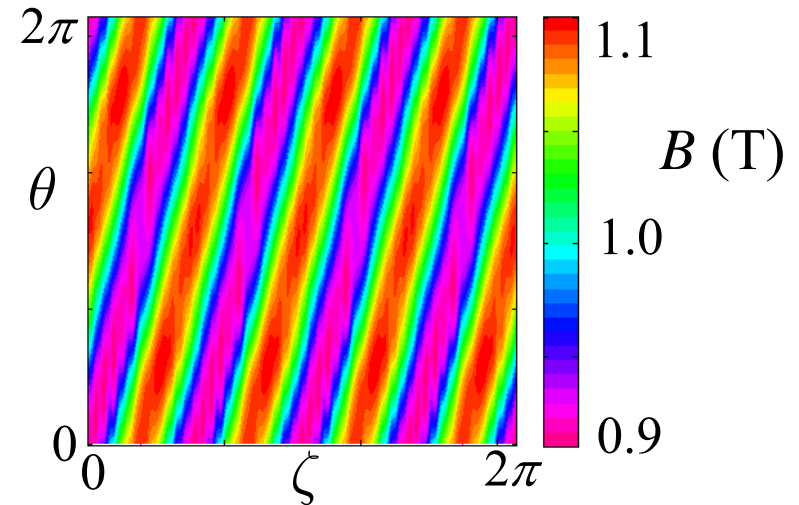
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## New geometric consequence of omnigenity:

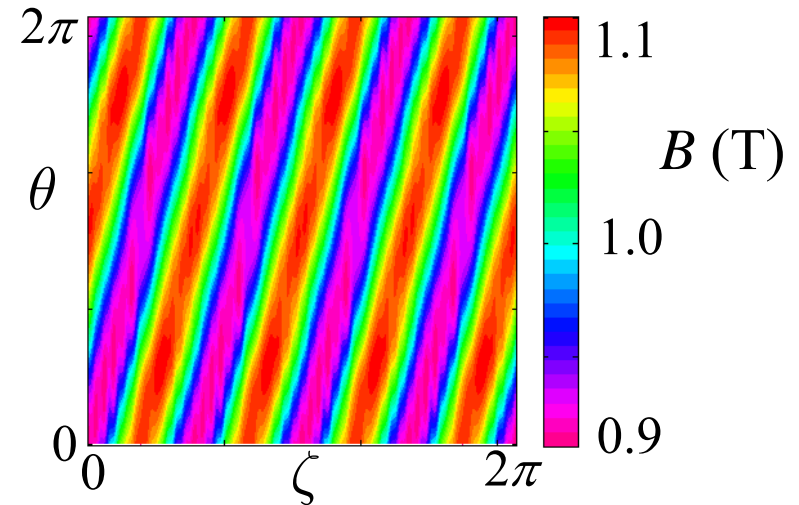
By applying Ampère's Law to a  
 $B$  contour on a flux surface,

$$\oint \mathbf{B} \cdot d\mathbf{r} = \frac{4\pi}{c} \times \underbrace{(\text{enclosed current})}_{MG + NI}$$

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$$\Rightarrow \frac{\mathbf{B} \times \nabla \psi \cdot \nabla B}{\mathbf{B} \cdot \nabla B} = \frac{2q}{c} \left( \frac{MG + NI + H}{M - qN} \right) \text{ where } \langle H \rangle = 0.$$

Current in an omnigenous plasma is described by a concise, explicit, analytical formula.

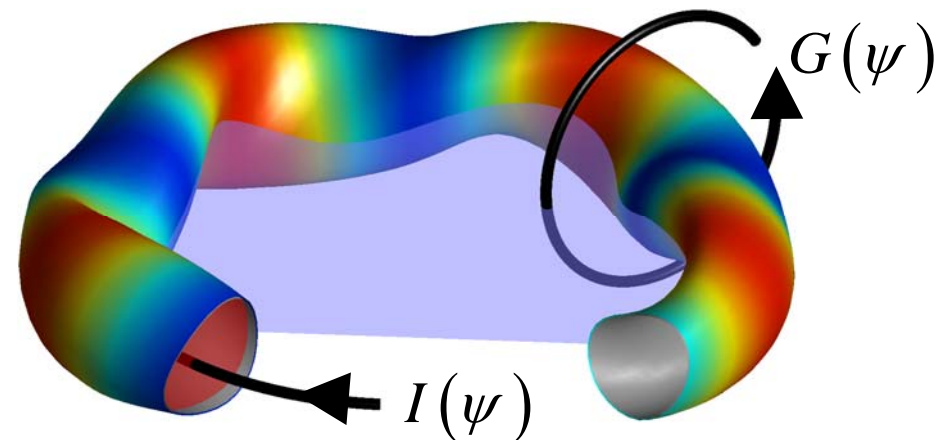
$$j_{\parallel} = 3.3 \frac{f_t q B}{\langle B^2 \rangle} \left( \frac{NI + MG}{qN - M} \right) \left( \frac{dp_e}{d\psi} + \frac{dp_i}{d\psi} - 0.74 n_e \frac{dT_e}{d\psi} - 1.17 n_e \frac{dT_i}{d\psi} \right) + \frac{2q}{B(qN - M)} \left( \frac{dp_e}{d\psi} + \frac{dp_i}{d\psi} \right) \left[ \left( 1 - \frac{B^2}{\langle B^2 \rangle} \right) (NI + MG) + W \right]$$

Tokamak result with  $G \rightarrow -(NI + MG) / (qN - M)$

$$W = \frac{2B^2}{q} (qG + I) \times \int_0^{\zeta} \frac{d\zeta'}{B'^3} \left( N \frac{\partial B'}{\partial \theta} + M \frac{\partial B'}{\partial \zeta} \right)$$

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$I(\psi)$  and  $G(\psi)$  are the toroidal & poloidal currents.



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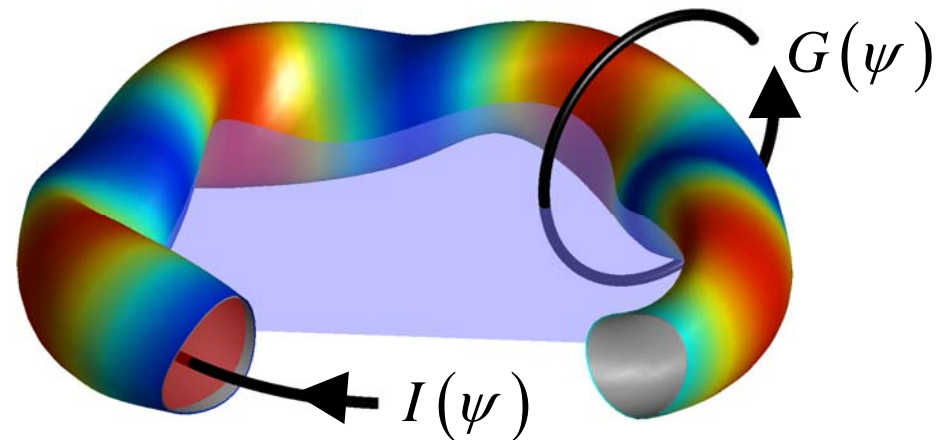
$$V_{\parallel i} = -1.17 \frac{2qB}{e \langle B^2 \rangle} \frac{dT_i}{d\psi} \frac{(NI + MG)}{(qN - M)} + \frac{2q}{B} \left( \frac{d\Phi}{d\psi} + \frac{1}{en} \frac{dp_i}{d\psi} \right) \frac{(NI + MG + W)}{(qN - M)}$$

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## Non-quasisymmetric stellarators:

- Neoclassical radial current depends on  $E_r$ .
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## Tokamaks & quasisymmetric stellarators:

- Radial fluxes of ions and electrons are always equal, regardless of  $E_r$  (“intrinsic ambipolarity”)

*(Helander & Simakov, PRL 2008)*

$\Rightarrow$  You **cannot** solve for  $E_r$ .



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$$\langle \mathbf{j} \cdot \nabla \psi \rangle = \left( Z e n_i \frac{d\Phi}{d\psi} + T_i \frac{dn_i}{d\psi} - 0.17 n_i \frac{dT_i}{d\psi} \right) \left\langle (\text{departure from quasisymmetry})^2 \right\rangle$$

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- $\langle \mathbf{j}_{\text{neoclassical}} \cdot \nabla \psi \rangle \gg \langle \mathbf{j}_{\text{turbulence}} \cdot \nabla \psi \rangle$ .

*(Helander & Simakov, Contrib. Plasma Phys. 2010)*

$\Rightarrow$  You can solve for  $E_r$  using  $\langle \mathbf{j}_{\text{neoclassical}} \cdot \nabla \psi \rangle = 0$ .

## Omnigenous stellarators:

$$\langle \mathbf{j} \cdot \nabla \psi \rangle = \left( Z e n_i \frac{d\Phi}{d\psi} + T_i \frac{dn_i}{d\psi} - 0.17 n_i \frac{dT_i}{d\psi} \right) \left\langle (\text{departure from quasisymmetry})^2 \right\rangle$$

Universal result:

$$\frac{d\Phi}{d\psi} = \frac{1}{Ze} \left( -\frac{T_i}{n_i} \frac{dn_i}{d\psi} + 0.17 \frac{dT_i}{d\psi} \right)$$

Totally independent of the details of  $\mathbf{B}$ .

# Summary: omnigenity is a useful limit.

- More general than quasisymmetry.
- Relevant (at least for insight and code benchmarking) to W7-X and to any reactor.
- Concise, explicit, analytical formulae for  $\mathbf{j}$ ,  $\mathbf{V}$ , and  $E_r$ .
- Poloidally closed  $B$  contours  $\Rightarrow$  zero bootstrap current.
- For non-quasisymmetric  $\mathbf{B}$ ,  $E_r$  is determined explicitly:

$$\frac{d\Phi}{d\psi} = \frac{1}{Ze} \left( -\frac{T_i}{n_i} \frac{dn_i}{d\psi} + 0.17 \frac{dT_i}{d\psi} \right) \quad \text{independent of the } \mathbf{B} \text{ geometry.}$$